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# Frontiers of Extreme Computing 2007 Applications and Algorithms Working Group

**October 25, 2007**

**Horst Simon (chair), David Bailey, Rupak Biswas, George Carr, Phil Jones,  
Bob Lucas, David Koester, Nathan Price, Joshua Schrier, Mark Stalzer (co-  
chair), David Womble**

# Outline

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- ◆ Sample Applications
- ◆ Exascale Applications Characteristics
- ◆ Zettascale Applications with Societal Impacts
  - Climate Modeling
  - Personalized Medicine and Phylogenomics
  - Fusion and Astrophysics
  - Decision Making
  - Nano/Material Science, Chemistry
- ◆ Adaptivity
- ◆ Requirements and Recommendations

# Sample Applications

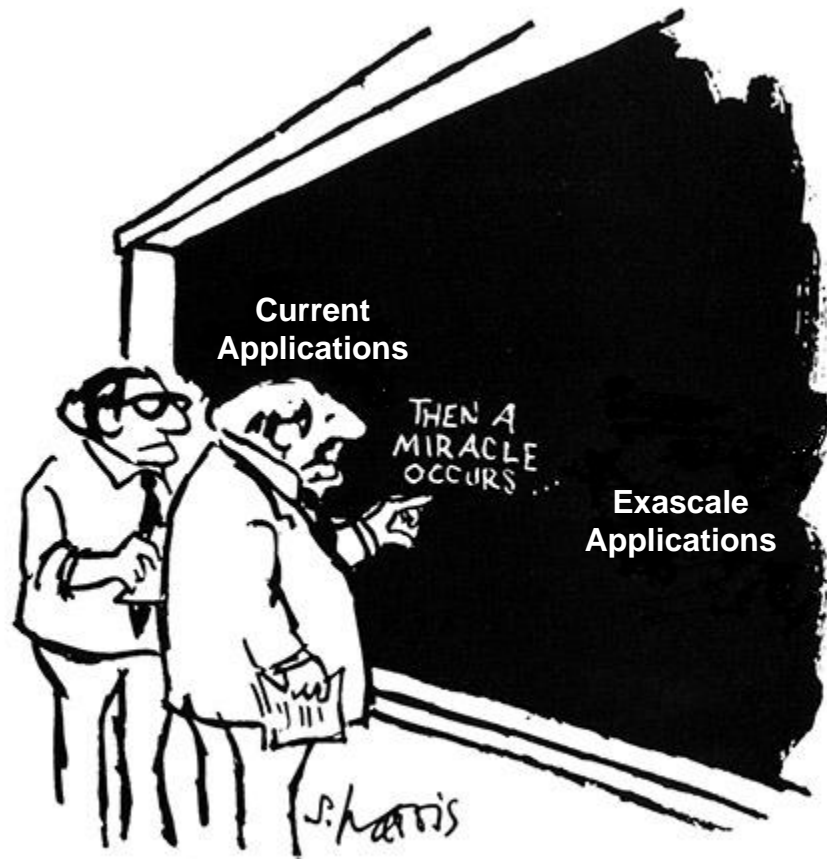
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- ◆ Climate modeling.
- ◆ Biology: spread/analysis of diseases; phylogenetics.
- ◆ Nano/materials: simulate from first principles a transistor/quantum computer.
- ◆ Chemistry: beyond Born-Oppenheimer.
- ◆ Economic/behavioral models: Baghdad in faster than real-time.
- ◆ Cognitive models of brain: vision; mapping out functions.
- ◆ Medical applications: personalized medicine.
- ◆ Astrophysics.

# Sample Applications, Cont.

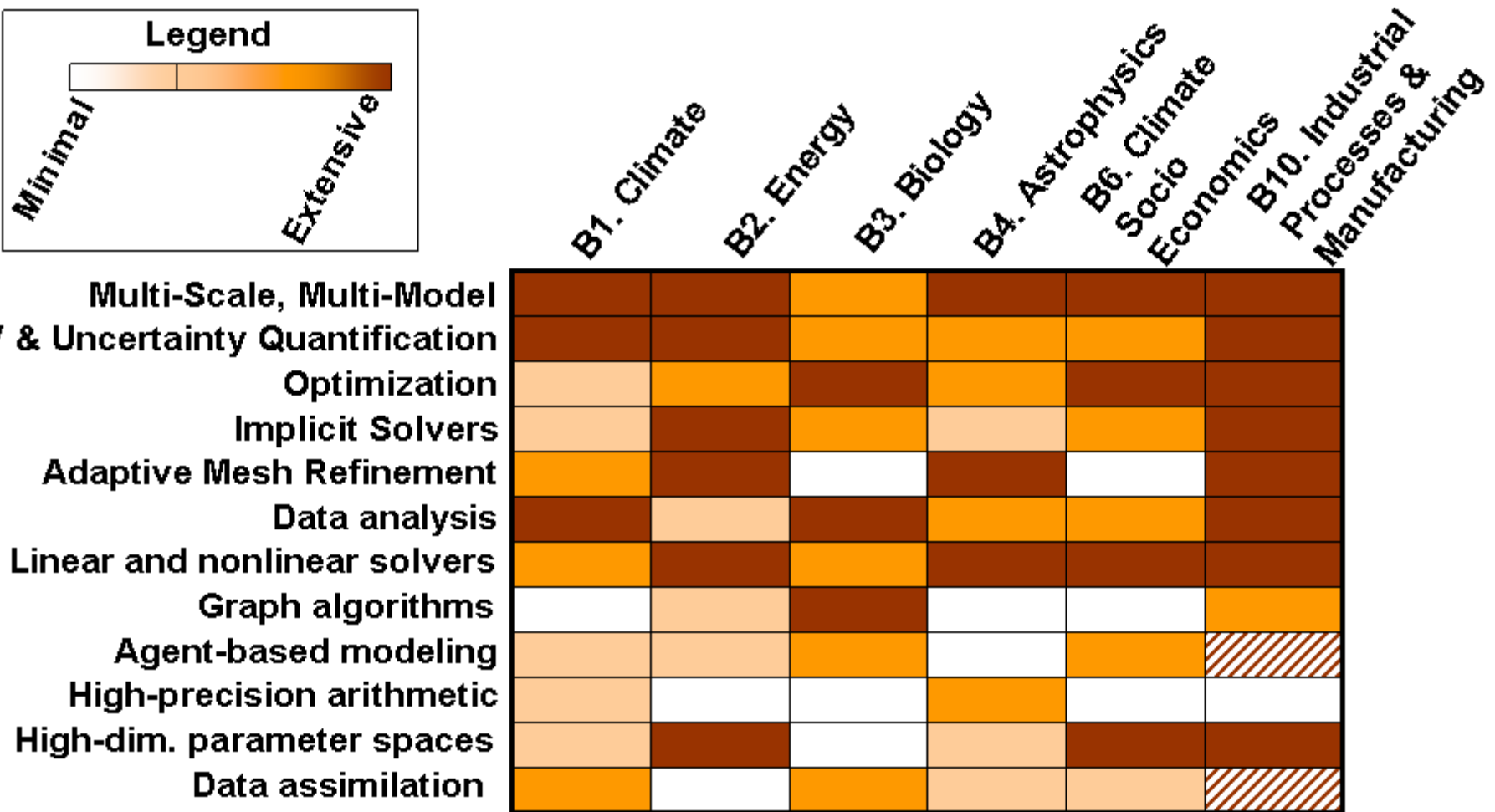
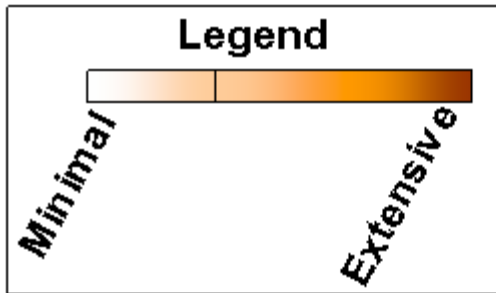
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- ◆ Engineering: crash modeling.
- ◆ Decision support.
- ◆ Calibration of large instruments (LHC, LIGO).
- ◆ HEP and nuclear physics.
- ◆ Real-time weather: storms, tornados, icing; disaster planning.
- ◆ Combustion.
- ◆ Nuclear energy: fission; fusion.
- ◆ Stockpile stewardship (ASC).
- ◆ Oil&gas exploration.
- ◆ Earth tomography.
- ◆ Math research.



"I think you should be more explicit here in step two."

# Exascale Application Characteristics



Application characteristics from Group B7 Mathematics and Algorithms, David Bailey, et.al.

# Climate Modeling

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- ◆ Application and impact on society
  - Provide projections of future climate to policy makers to support decisions on energy, mitigation and adaptation strategies.
  - Require accurate models of both physical and biogeochemical climate system.
  - Require projections at regional spatial scales.
- ◆ Algorithms
  - Coupled multi-model application.
  - Wide variety of algorithms.
  - Largely explicit forward integration of PDEs.
  - Future needs: implicit time, efficient methods for large tracer counts, more efficient strategies for grids, discretizations and decomposition.

# Computing Needs (10<sup>10</sup>-10<sup>12</sup>)

## Resolution (10<sup>3</sup>-10<sup>5</sup>)

- x10 each x&y, x10 timestep, x5-10 z
- Regional prediction (10km)
- Eddy resolving ocean (10km)

## Completeness (10<sup>2</sup>)

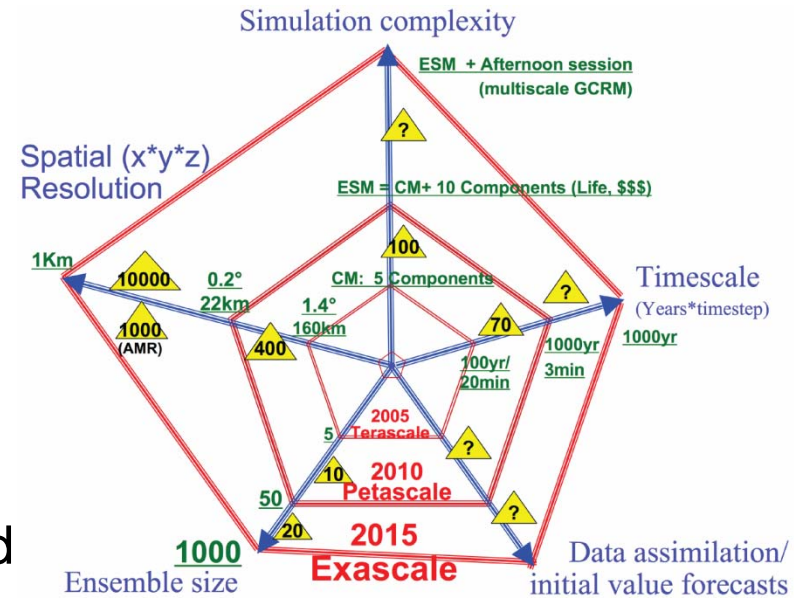
- Biogeochem (30-100 tracers, interactions)
- Ice sheets

## Fidelity (10<sup>2</sup>)

- Better cloud processes, dynamic land etc.

## Increase length/number of ensembles (10<sup>3</sup>)

- Run length (x100)
- Number of scenarios/ensembles (x10)
- Data assimilation (3-10x)





# Storage/Data Requirements

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- ◆ Currently 10 GB/year output
  - Similar multipliers (min x1000).
  - Staged migration to archival systems.
  - 2x to move initially, recall many times.
  - At 24 years/day => 10 GB/hour I/O rates.
  - Much higher data input rates if data assimilation required.
- ◆ Analysis
  - Requires large memory systems.
  - Need lots of intermediate disk (don't want to analyze from archival systems).
  - Distributed data (IPCC multi-model analyses).

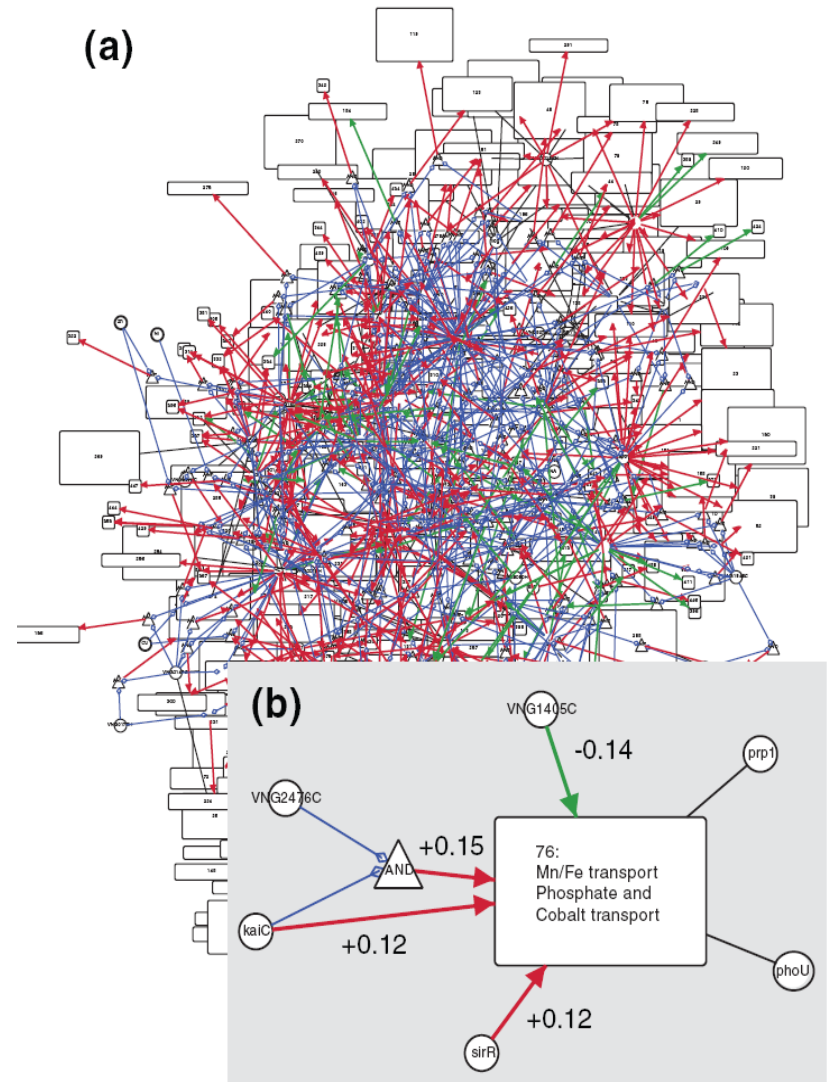
# Personalized Medicine

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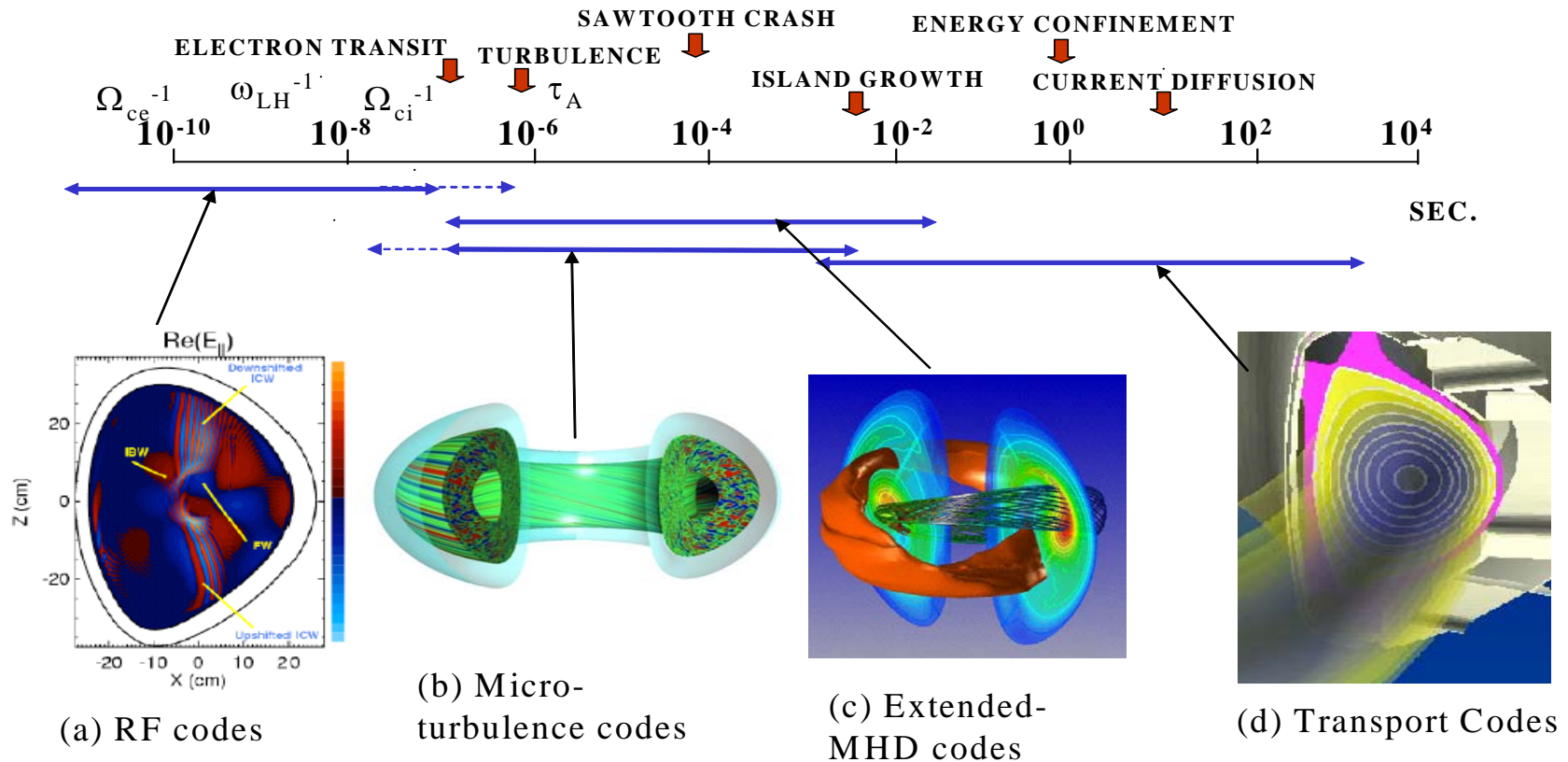
- ◆ Personalized Medicine - Identifying genomic predictors of disease from individual human genomes.
- ◆ Resources: Possible sets of mutation sets of size 6 (e.g. typical for cancer) in the human genome  $\sim 10^{55}$ .
  - Thus, we will need significant computational resources and efficient approximation algorithms to best deliver on personalized medicine (e.g. for cancer).
  - For even a simple binary classification, would thus need  $\gg 185$  individual genomes to begin to have hopes of separating signal from noise – i.e. to believe good patterns that are found (i.e.  $10^{55} \sim 2^{185}$ ) (and this assumes simplistic yes/no model selection).
  - But there is also a decreasing probability of finding good patterns as number of classifications is done.

# Phylogenomics

- ◆ Phylogenomics - learning about the origin and history of life.
- ◆ Resources: Phylogenomic calculations are NP-complete; approximate solutions get better with increased computational power, as more accurate but computationally intensive algorithms are used  $O(N, N \log N, N^2, \dots)$ .



# Fusion Code Characteristics



- ◆ **Multi-physics, multi-scale computations.**
- ◆ **Numerous algorithms and data structures.**
- ◆ **Regular and irregular access computations.**
- ◆ **Adaptive mesh refinement.**
- ◆ **Advanced nonlinear solvers for stiff PDEs.**

# Fusion Requirements

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Tokamak turbulence (GTC) - For ITER experiment, etc:

- Grid size: 10,000 x 4000 x 256, or about  $10^{10}$  gridpoints.
- Each grid cell contains 8 particles, for total of  $8 \times 10^{10}$ .
- 200,000 time steps required.
- Improved plasma model (increases by 10-100X).
- Total cost:  $6 \times 10^{21}$  flop = 1 hours on 1 Eflop/s system; 10 Pbyte main memory.

All-Orders Spectral Algorithm (AORSA) - to address absorption of RF electromagnetic waves in plasmas.

Present Day:

- 120,000 x 120,000 complex linear system requires 230 Gbyte memory, 1.3 hours on 1 Tflop/s.
- 300,000 x 300,000 linear system requires 8 hours.

Future (ITER scale):

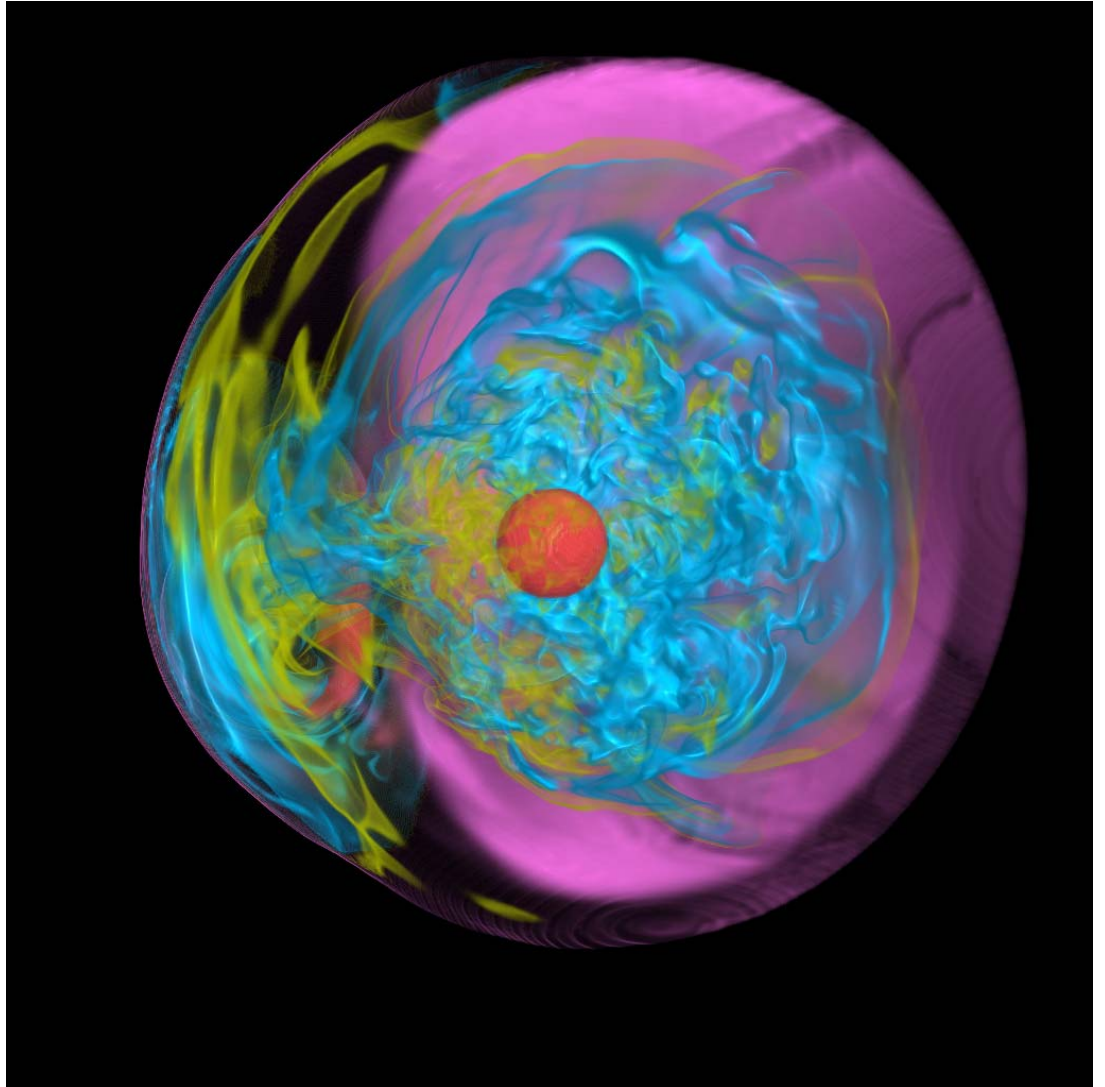
- 6,000,000 x 6,000,000 system = 1 hour on 1 Eflop/s system; 1 Pbyte memory.

# Supernova Physics

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Supernova shock  
wave instability from  
3D core collapse  
supernova simulation.

[K.L. Ma (viz) and  
J.M. Blondin]



# Astrophysics Code Characteristics

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- ◆ Undergoing a transformation from a data-starved discipline to a data-swamped discipline.
- ◆ Large effort in experimental data analysis (microwave background and supernova)
- ◆ Typical simulation applications: supernova hydrodynamics, energy transport, black hole simulations.
- ◆ Multi-physics and multi-scale phenomena.
- ◆ Large dynamic range in time and length.
- ◆ Requires adaptive mesh refinement.
- ◆ Dense linear algebra.
- ◆ FFTs and spherical harmonic transforms.
- ◆ Operator splitting methods.

# Astrophysics Requirements

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Supernova simulation:

- 3-D model calculations will require 1M processor-hours per run, on 1 Pflop/s system, or 1000 hours per run on 1 Eflop/s system.

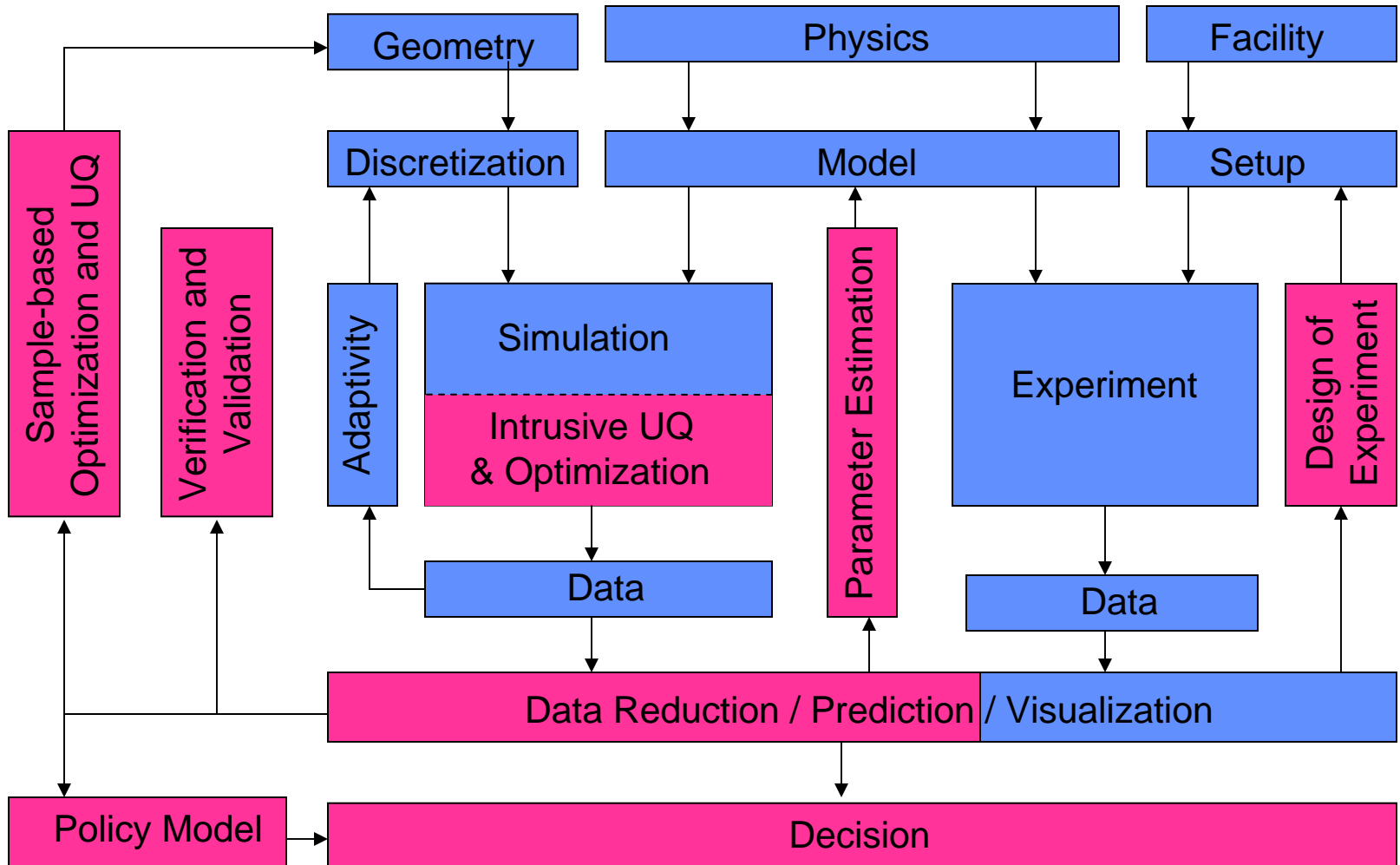
Analysis of cosmic microwave background data:

- WMAP (now)                       $3 \times 10^{21}$  flops, 16 Tbyte mem
- PLANCK (2007)                 $2 \times 10^{24}$  flops, 1.6 Pbyte mem
- CMBpol (2015)                 $1 \times 10^{27}$  flops, 1 Ebyte mem

Note: Microwave background data analysis, and also supernova data analysis, involves mountains of experimental data, not simulation data.



# Decision Making Flow



# Decision Making

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- ◆ Any code can be put into a decision making framework to ask
  - Do I believe the model/code? (i.e., V&V)
  - How good? What if? How can I make it better? How can I achieve this through policy? (i.e., design)
- ◆ Needs
  - Optimization and uncertainty quantification (includes model building, e.g., setting knobs).
    - sampling-based (need 100x or more computation, capacity computing okay).
    - Intrusive (need 10x computation, need same capability as the application).
  - Policy modeling
    - Looks like discrete event simulation.
    - Based on communication, not computation (need low latency communication and global synchronizations).
  - Data reduction / Prediction / Coupling to experiment
    - Large data sets – need large-scale storage.
    - Streaming data – need bandwidth and multiple threads.

# Nano/Material Science, Chemistry

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- ◆ Objective:
  - *Ab initio* million-atom electronic structure simulations.
- ◆ Impact:
  - Atomistic simulation and design of nanoscale-, molecular-, quantum-information processing devices  $10^6$  atoms, 1-10 ns timescales.
- ◆ Resources (time/mem/disk):
  - Local DFT:  $N-N^2-N^3$  /  $N-N^2$  /  $N-N^2$ 
    - Current state of the art:  $O(10^3)$  atom static,  $O(50)$  atoms dynamics.
    - QBOX: 207 Tflop/s.
    - Strategies to avoid orbital basis exist, still require science+algorithms.
  - Hybrid DFT:  $N-N^4$  time (e.g., Link)
- ◆ Driver for computer architecture:
  - FFT and associated communication bottleneck.
    - Primarily an issue in planewave methods, however, utilized in some  $O(N)$  localized orbital methods (but currently small prefactor).
  - Both dense AND sparse linear algebra problems.
  - Global arrays successful in NWChem.

# Adaptivity

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- ◆ Adaptivity will be required even in zettaflops regime
  - Some problems may not require more resolution beyond a certain point, but adaptivity will improve overall efficiency and time-to-solution.
  - Other problems will benefit from localized higher resolution to capture physics, chemistry, other characteristics of the problem.
  - Minimal floating-point ops, mostly conditionals and branching.
  - Merely a tool to increase efficiency and not a goal in itself.
- ◆ Adaptivity of two types:
  - hpr-refinement (structured, unstructured, Cartesian grids, overset grids).
  - Domain specific (e.g. add appropriate models at certain resolutions).
- ◆ Adaptivity can be done in parallel (has sufficient concurrency), but:
  - Will need synchronization at processor / domain boundaries unless requirements relaxed from mesh consistency and numerical accuracy.
  - Mesh and solution quality issues become more important in zettaflops regime because # of refinement levels expected to increase exponentially.

# Adaptivity, Cont.

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- ◆ Multi-objective dynamic load balancing needed for adaptivity
  - Balance computations (multi-disciplinary, multi-science, multi-physics).
  - Balance communications (inter-processor, processor-to-memory).
  - Will have to do load balancing at a higher level (since  $\#DOF \gg \#processing\ units$ ).
- ◆ Solution quality
  - Estimating error in the solution generally more expensive than generating solution.
  - Must have built-in sensitivity / uncertainty analyses to generate confidence in numerical solutions.
  - Holy grail is obtaining grid-independent solutions.

# Exascale Performance Requirements

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A balanced exaflops machine needs:

- ◆ Main memory: ~400 PB
- ◆ Working storage: 4 to 40 EB
- ◆ Archival storage: 40 to 1000 EB
- ◆ Bisection Bandwidth: 0.5 to 1 EB/s
  - Climate pushes local bandwidth
  - FFT pushes global bandwidth
- ◆ Balanced RandomAccess and HPL performance

# Other Exascale Requirements

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- ◆ Support legacy applications:
  - \$2.5B (just DOE & NSF) MPI software base.
  - 20 year lead time for new programming models.
- ◆ Develop new petascale programming environments with exascale in mind:
  - A mere 100,000 threads is difficult too!
  - New algorithms opportunity to use new programming models.
- ◆ Fault tolerance (real system throughput):
  - Error checking and notification.
  - System level checkpoint/restart.
- ◆ Chip & network performance monitoring

# Other Issues

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- ◆ Extensions to Von Neumann model?
  - Better support for adaptive codes
  - Rick Stevens applications
- ◆ Seamless integration with distributed data-rich computing environment.